

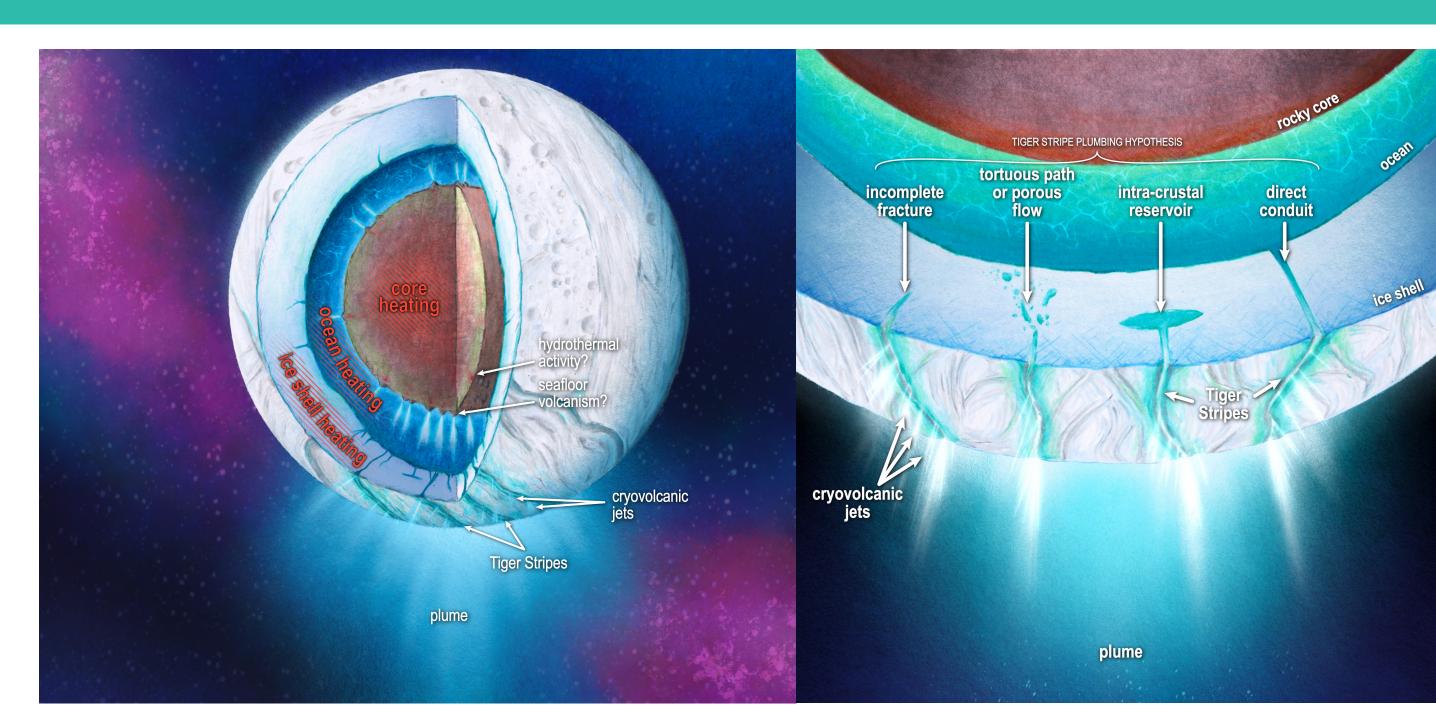
FY23 Strategic Initiatives Research and Technology Development (SRTD)

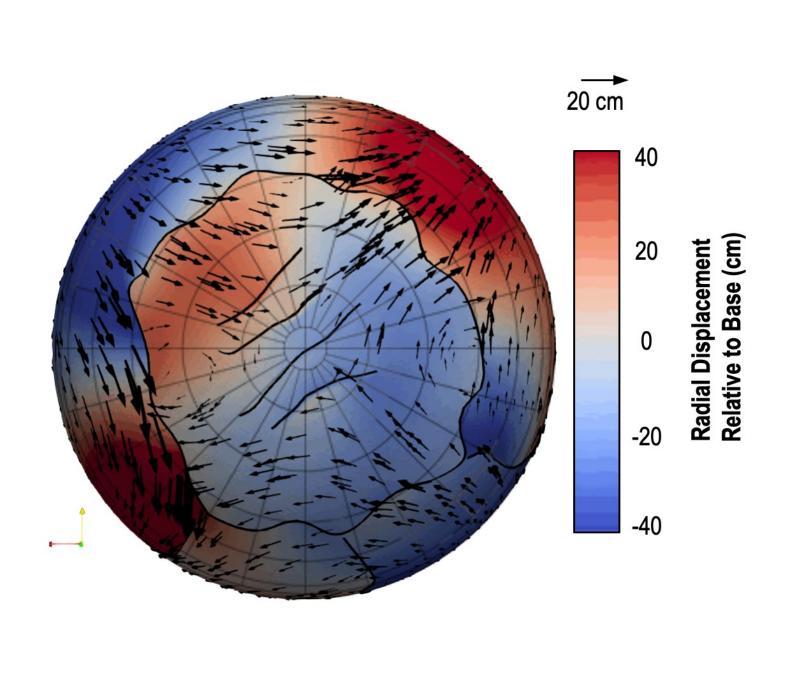
## Determining the Scientific Impact of a Geodesy Network at Enceladus

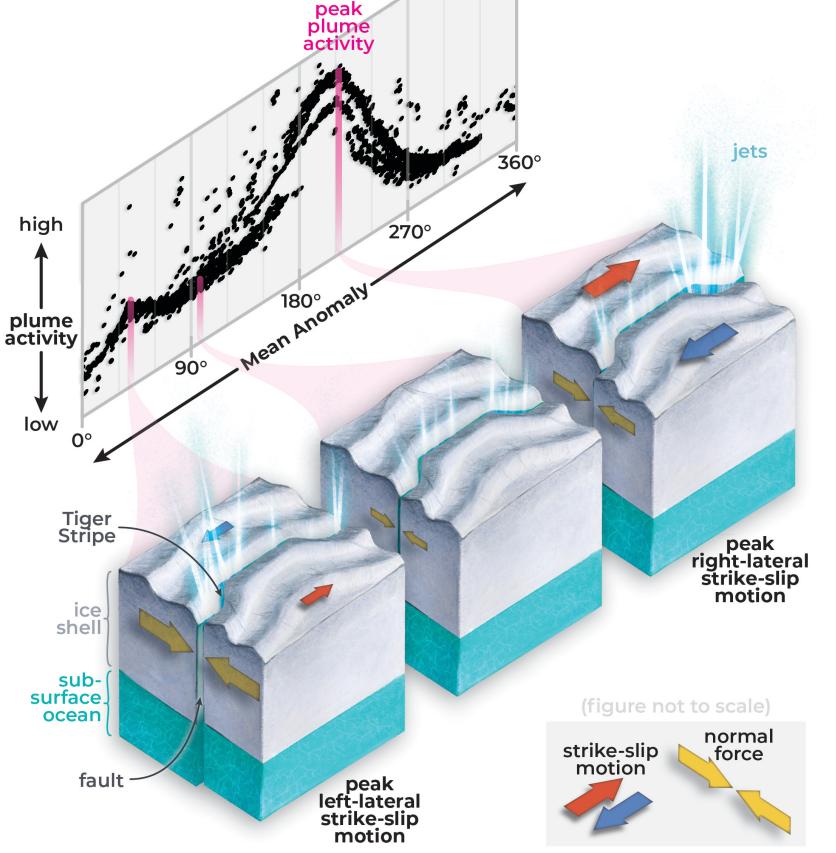
**Principal Investigator:** James Keane (322); **Co-Investigators:** Ryan Park (392), Steven Vance (322), Paul Rosen (330), Mark Simons (Caltech), Alex Berne (Caltech)

Strategic Focus Area: Next-Generation Ocean World Geodesy: Enceladus | Strategic Initiative Leader: Steve Vance

**OBJECTIVES:** The overarching goal of this R&TD is to develop the fundamental scientific tools that tie to science questions to capabilities that could be provided by next-generation ocean world geodesy, with a focus on Enceladus. At present, there are critical gaps in available geophysical models of Enceladus, inhibiting the community's ability to testable hypotheses, with quantitative predictions for measurements, and ultimately answer science questions. The primary objective for this year was to develop a new global geophysical model of Enceladus capable of simulating crustal deformation which incorporates viscoelastic rheology, shell thickness variations, faulting, and other relevant processes at high spatial and temporal scale. Most state-of-theart geophysical models of Enceladus are built on questionable simplifying assumptions, such as the treatment of faulting, assumed rheological models, the role of faulting on global deformation, and global symmetry. Enceladus's faults are the conduits by which ocean material is ejected into space, yet we do not understand how they form or evolve, nor their role in Enceladus's past/present habitability. We have developed a new finite element model of Enceladus using PyLith—a toolkit developed and rigorously tested for Earth applications.







**APPROACH & RESULTS:** We have developed a new global geophysical model of Enceladus, capable of simulating realistic crustal deformation by incorporating ice shell thickness variations, faulting, viscoelastic rheology, and other relevant processes—all modeled at high spatial and temporal scale. This model, called SatDef, is based on the 3D viscoelasto-plastic finite element code, PyLith, which is well established tin the terrestrial geoscience community. We have benchmarked SatDef against analytical models (e.g., SatStressGUI) and previously published finite element models (e.g., Běhounková et al. 2017). We are continuing to add capabilities to SatDef, and are now entering the phase of the R&TD where we can explore the parameter space and craft testable hypotheses that are the necessary for future geodetic investigations of Enceladus. One of the most recent, exciting results is that we are finding that plume output correlates well with modeled strike-slip motion.

SIGNIFICANCE & BENEFITS TO JPL/NASA: This R&TD is directly relevant to JPL's strategic goals of exploring ocean worlds like Enceladus, and investigating their habitability. Geodesy provides one of the few ways to interrogate the hidden interiors of ocean worlds at a range of spatial and temporal scales—synergistic with seismology and electromagnetic methods. This R&TD is particularly timely, as it will enable JPL to propose competitive Discovery and New Frontiers missions that explore habitability with a multi-pronged, wholistic approach—complementing and enhancing life-detection investigations. This work can feed forward to the Enceladus Orbilander flagship in the future, and have applications for a variety of other worlds (e.g., Io, Europa, etc.). This work supports JPL's long-standing history of leadership in planetary geophysics.

## **National Aeronautics and Space Administration**

Jet Propulsion Laboratory
California Institute of Technology
Pasadena, California

www.nasa.gov

Clearance Number: CL#00-0000 Poster Number: RPC# Copyright 2023. All rights reserved.

## **Publications:**

 Berne, A., M. Simons, J. T. Keane, and R. S. Park. 2023. Inferring the Mean Thickness of the Outer Ice Shell of Enceladus From Diurnal Crustal Deformation. Journal of Geophysical Research (Planets) 128: e2022JE007712. DOI: 10.1029/2022JE007712.

## PI/Task Mgr. Contact Information:

James Tuttle Keane (james.t.keane@jpl.nasa.gov)